

HERMES: a constellation of nano-satellites for
high energy astrophysics and fundamental
physics research

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Lessons learned

Vela satellites, IPN, BeppoSAX, Swift

Distributed instrument → arcmin-deg positions

Modularity → improved performances

Prompt arcmin-arcsec positions → game changer

Mission concept

Disruptive technologies: cheap, underperforming, but producing high impact. Distributed instrument, tens/hundreds of simple units

HERMES constellation of cubesat

2016: ASI funds for detector R&D

2018: MIUR funds for pathfinder
(Progetti premiali 2015)

2018 H2020 Space-SCI-20 project

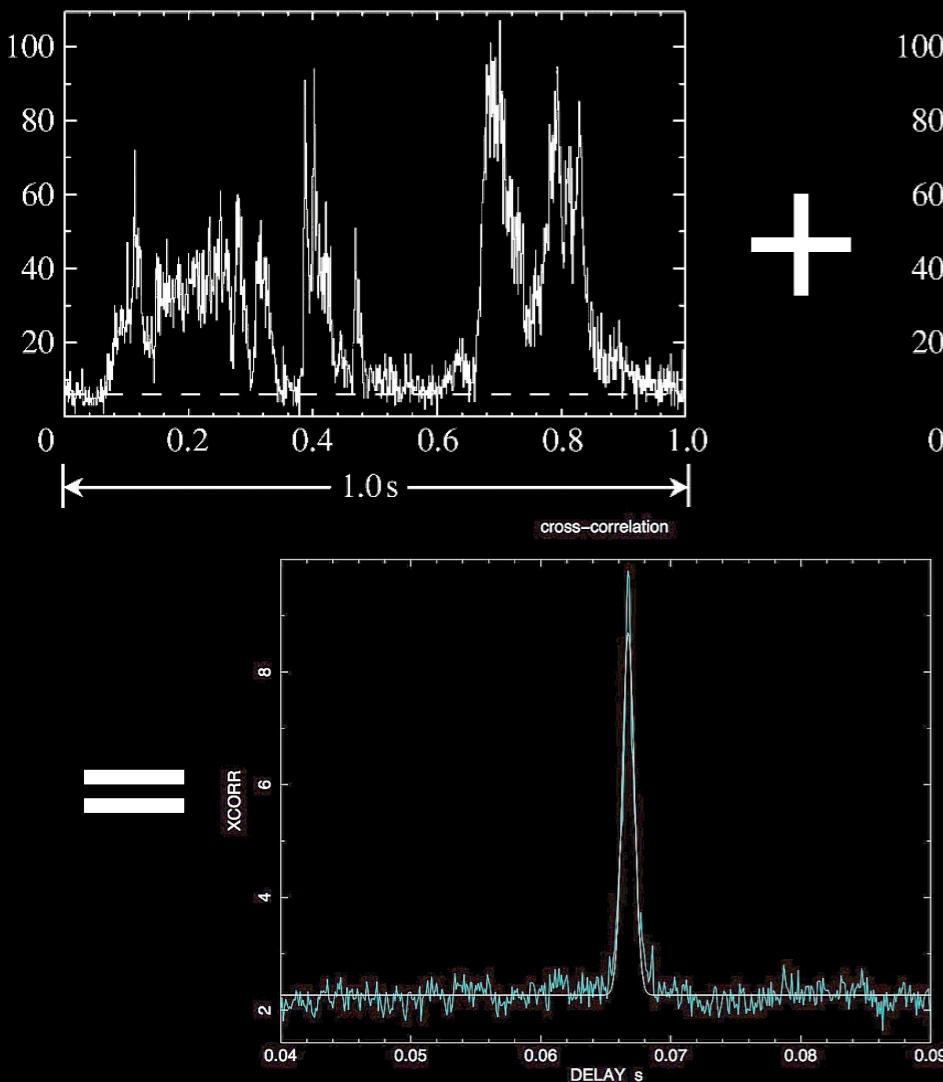
2018 ASI internal proposal



Experiment concept

1. Measure GRB positions through delays between photons arrival times:

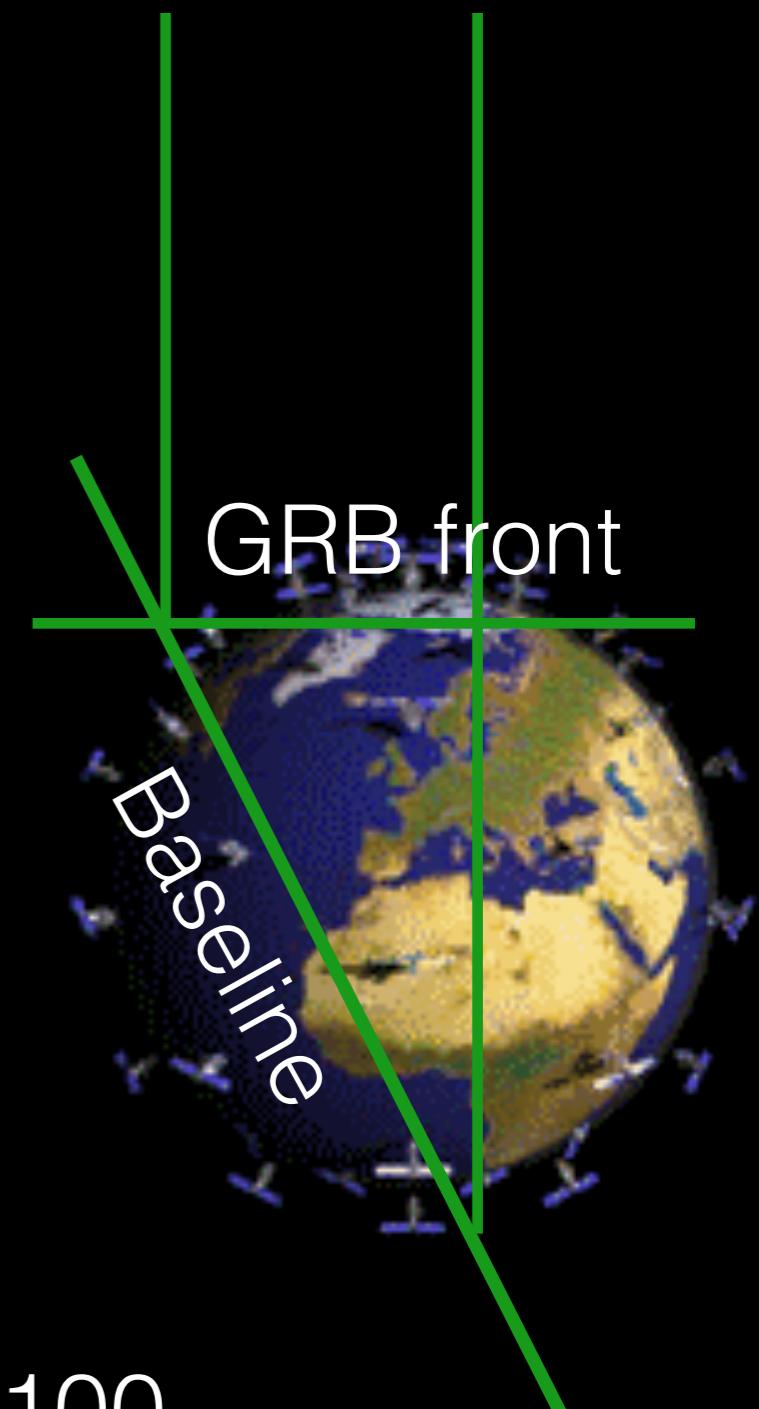
$$\sigma_{\text{Pos}} = \sigma_{\text{CCF}} \times c / \langle B \rangle / (N \times (N - 1) / 2)^{1/2}$$



$\sigma_{\text{CCF}} \sim 10 \mu\text{s}$

$\sigma_{\text{Pos}} \sim 10 \text{ arcsec}$

if $\langle B \rangle \sim 7000 \text{ km}$, $N \sim 100$



Experiment concept

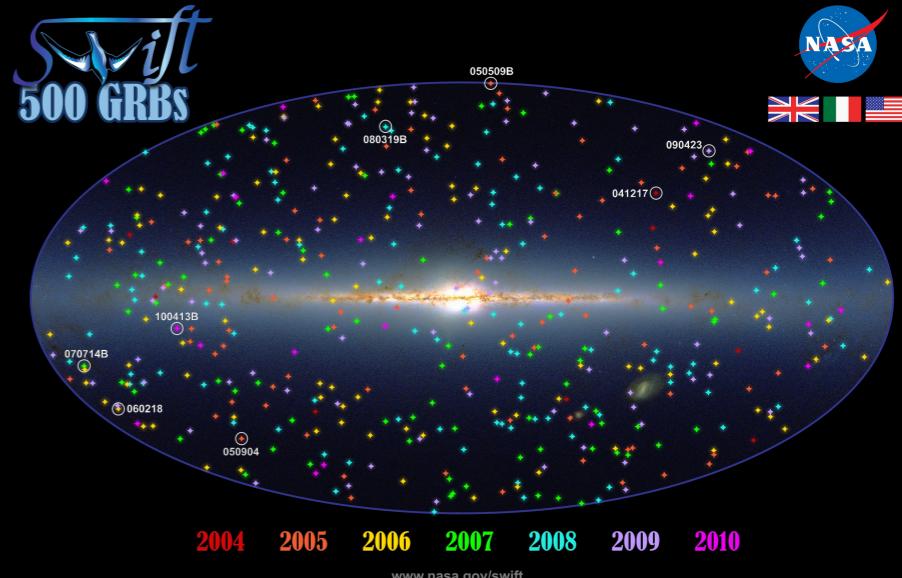
2. Add the signal from different units

Total collecting area 50-100-
 $\text{cm}^2 \times 100-200 = 0.5-2 \text{ m}^2$

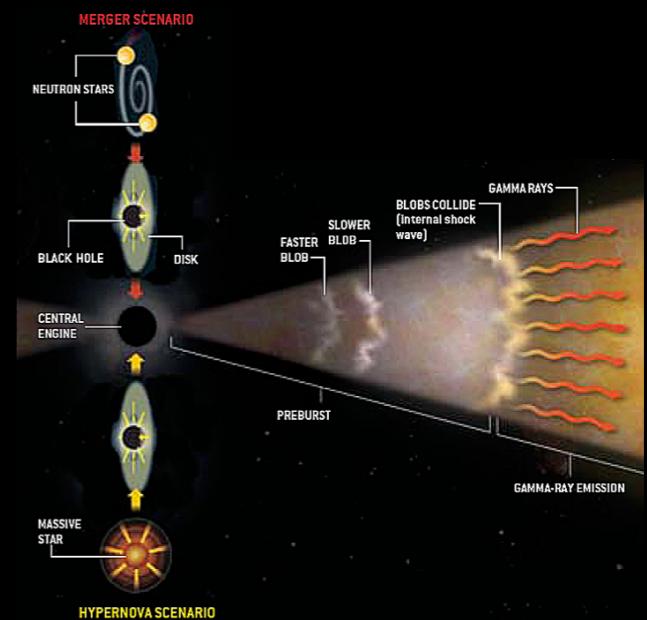
Transient fine ($\mu\text{s-ms}$)
temporal structure



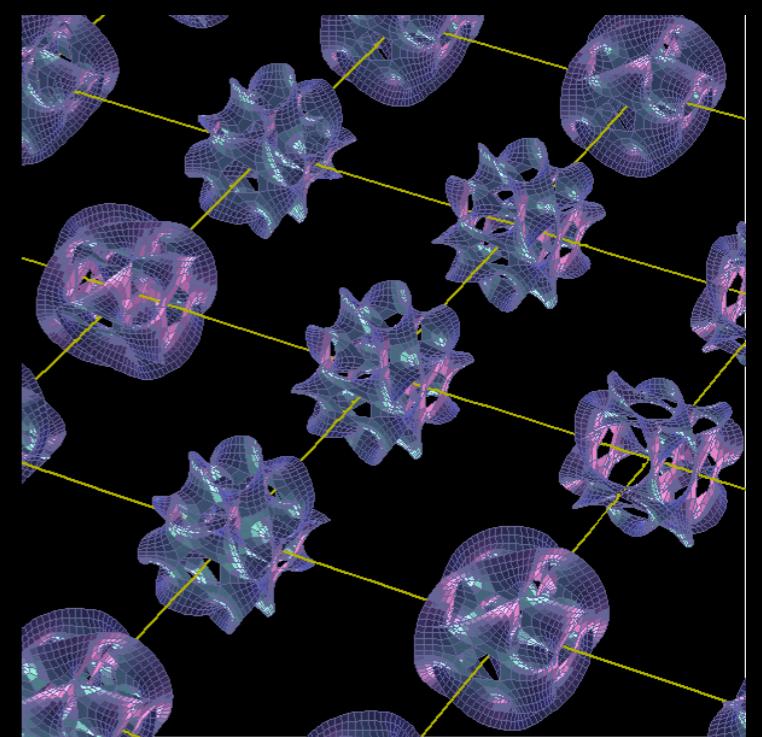
How to *promptly* localise a GRB *prompt* event?



How to construct a GRB engine?



Which is the ultimate granular structure of space-time?



Requirements

Scientific:

Arcmin-arcsec positions of ~a few dozen GRB/yr

Prompt(minute) localisation

μ s timing

$\Delta t / \Delta E \sim 3\mu s / 100 keV$ $30\mu s / 1 MeV \rightarrow M_{QG} \sim M_{Planck}$

Requirements

System:

≈hundreds detectors

single collecting area $\geq 50\text{cm}^2$

total collecting area $\geq 1\text{m}^2$

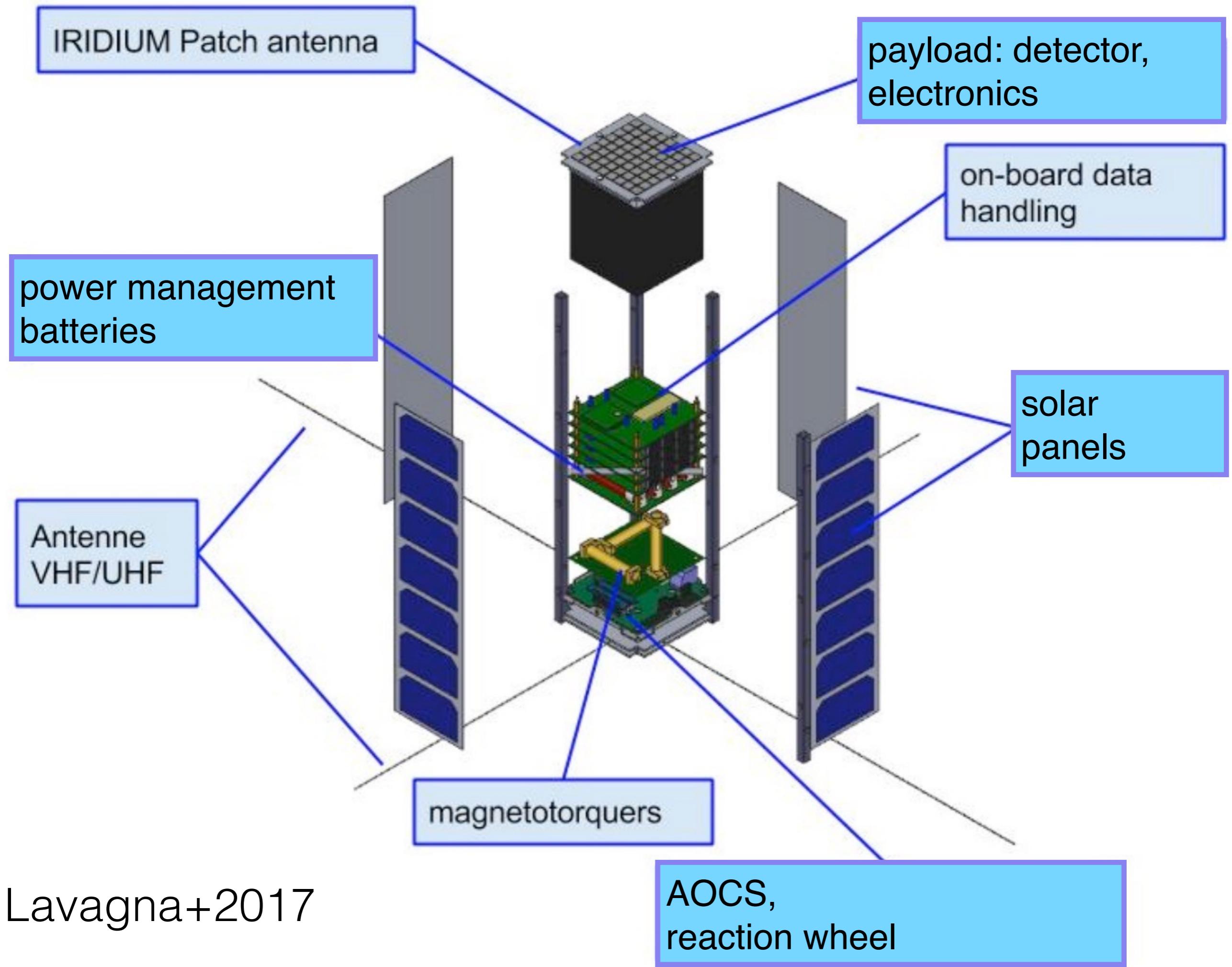
Energy range 3-10 — 300-1000 keV

Temporal resolution 10-100ns

Position reconstruction of each satellite < a few m

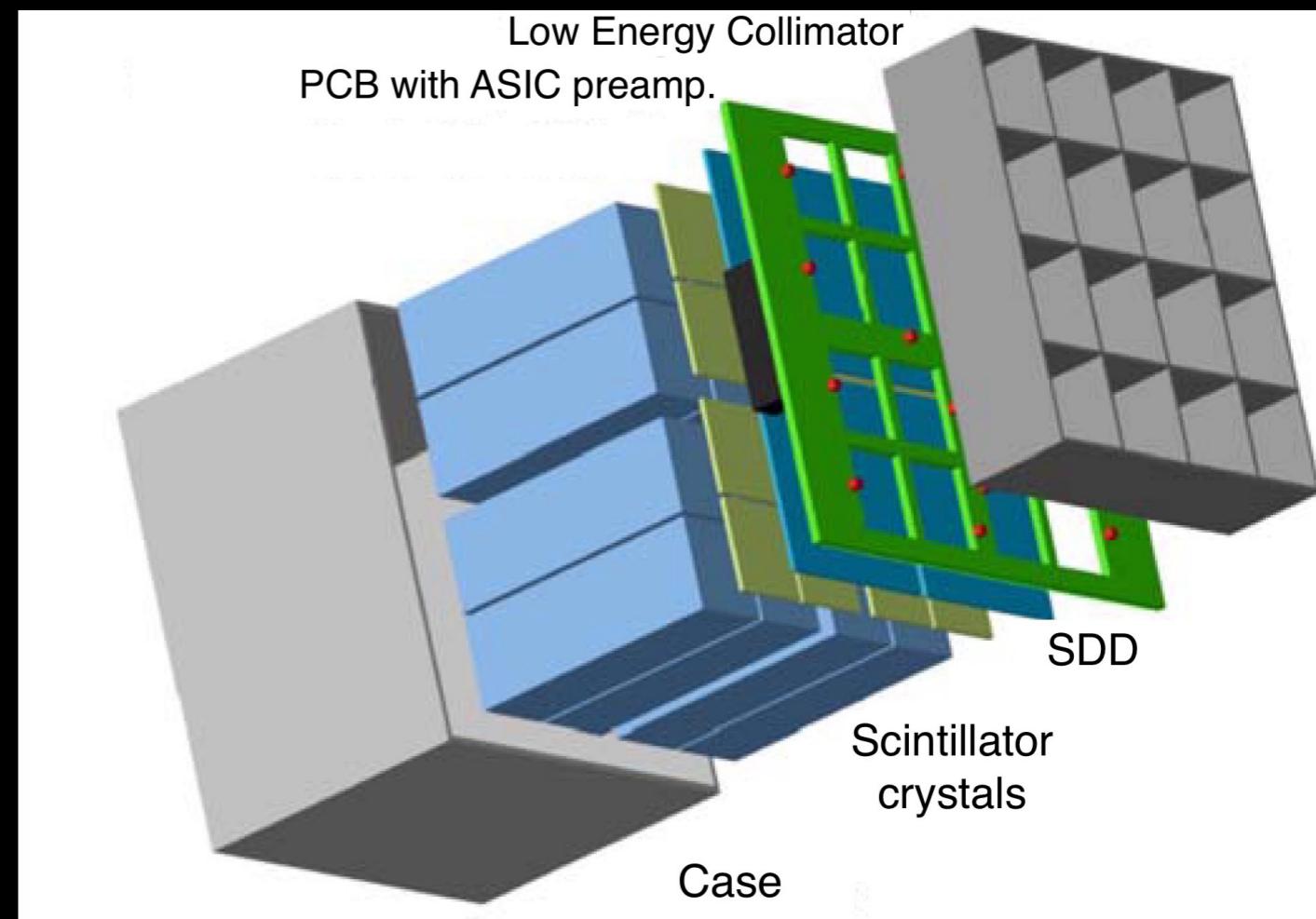
Absolute time reconstruction <10-100 ns

Download full burst info in minutes



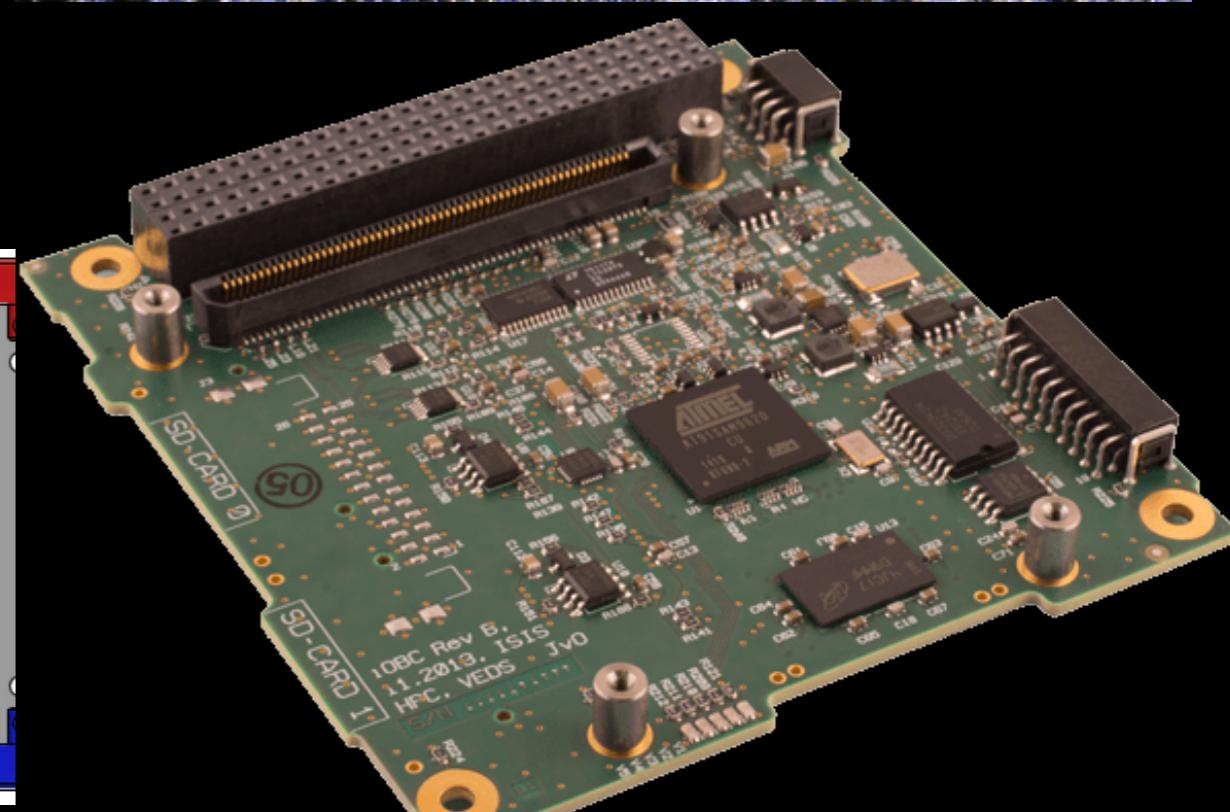
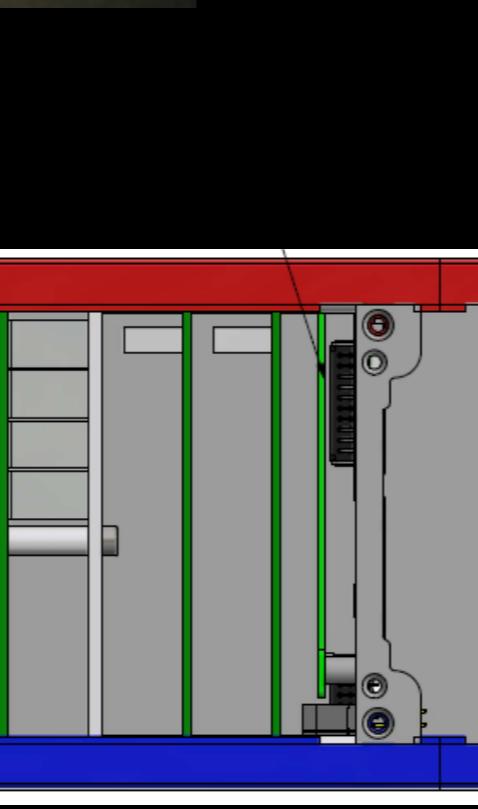
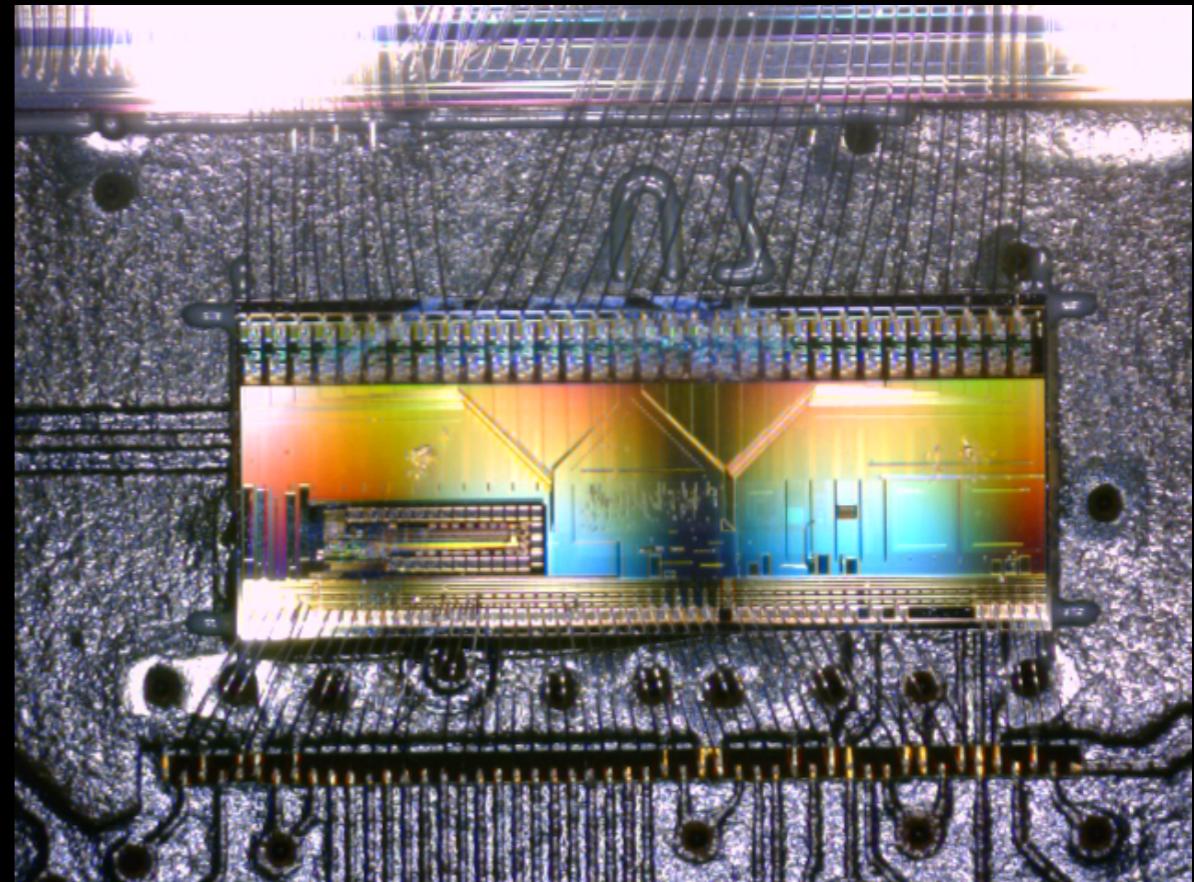
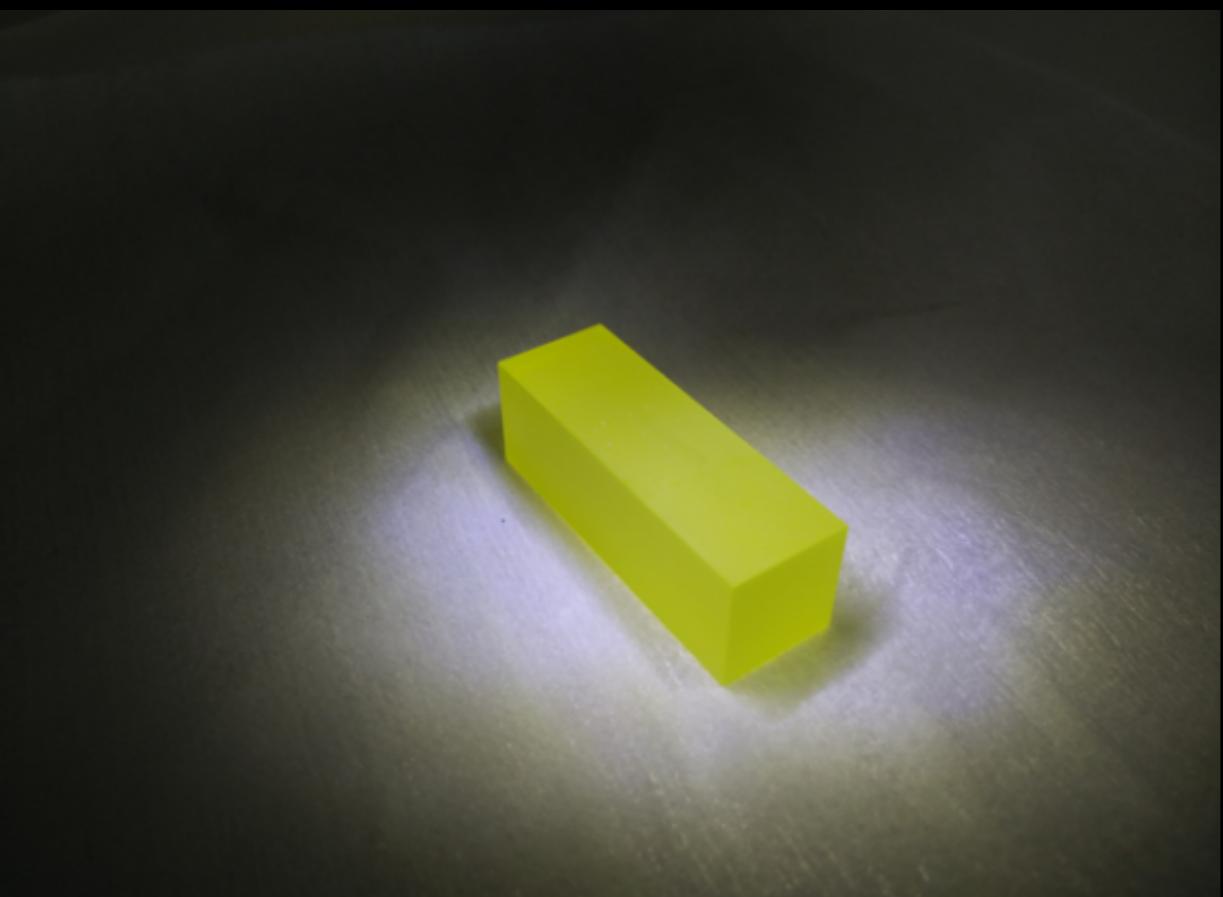
Payload

- Scintillator cristal GAGG
Photo detector, SDD
- 5-300 keV (3-1000 keV)
- ~50 cm² coll. area
- a few st FOV
- Temporal res. 10-100 nsec
- ~1.8kg



Fuschino+2018
Evangelista+2018
Campana+2018

From ppt to CAD to real stuff...



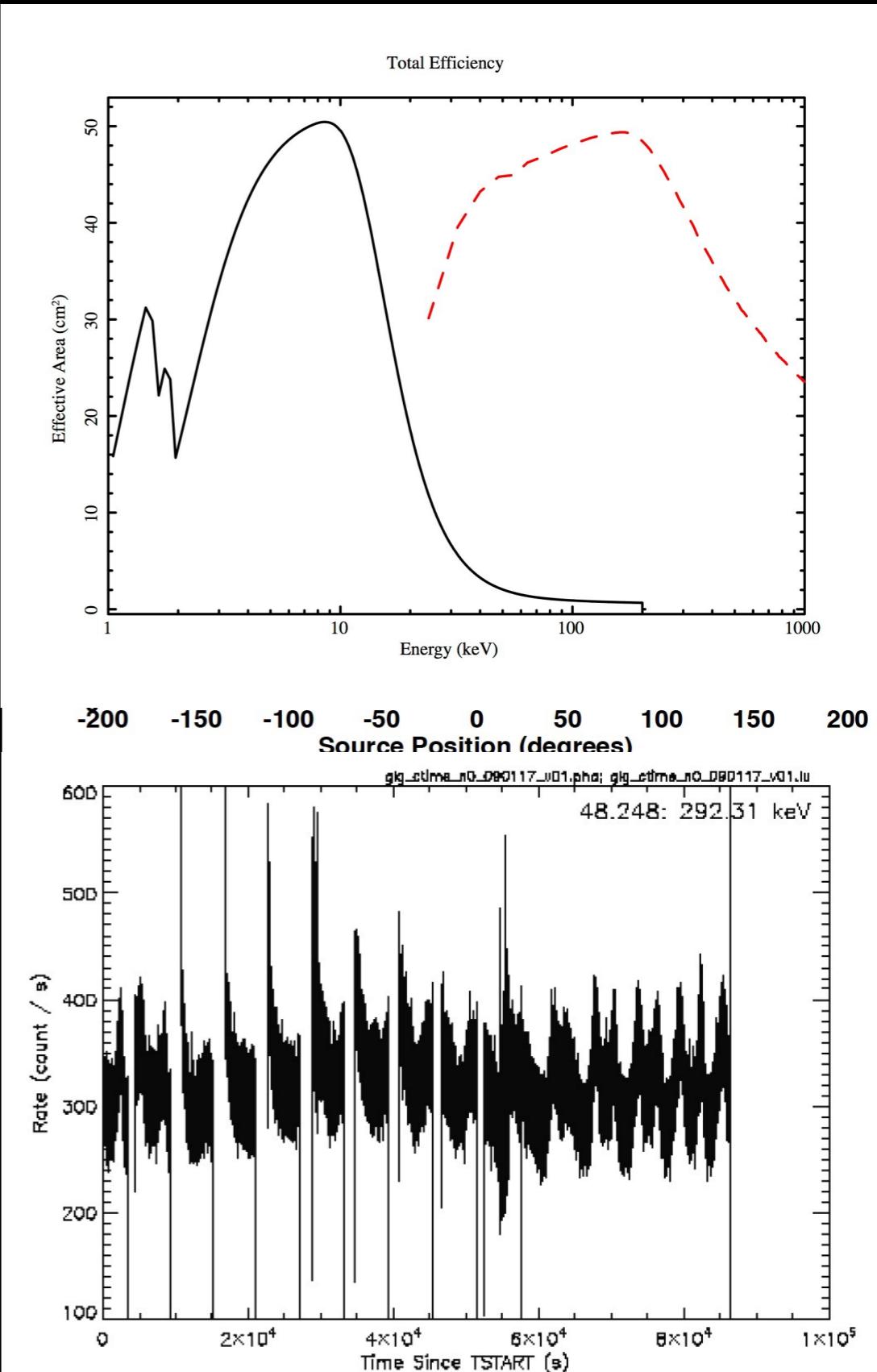
HERMES performances

Assumptions:

Instrument ~ 1 GBM module
~ 100cm^2 collecting area

Offaxis response ~ a few sterad
Background ~ 300-500 cts/s
50-300 keV

Minimum detectable count rate
~ 1 ph/ cm^2/s



HERMES performances

$$\sigma_{\text{Pos}} = 2.4^\circ [(\sigma_{\text{CCF}}^2 + \sigma_{\text{sys}}^2)/(N-3)]^{0.5}$$

$\langle B \rangle \sim 7000 \text{ km}$

$N(\text{pathfinder}) \sim 6-8$, active simultaneously 4-6

$N(\text{final constellation}) \sim 100$, active 50

$\sigma_{\text{Pos(pathfinder)}} \sim 1 \text{ arcmin}$ if $\sigma_{\text{CCF}}, \sigma_{\text{sys}} \sim 10 \mu\text{sec}$

$\sigma_{\text{Pos(FC)}} < 1 \text{ arcsec}$ if $\sigma_{\text{CCF}}, \sigma_{\text{sys}} \sim 10 \mu\text{sec}$

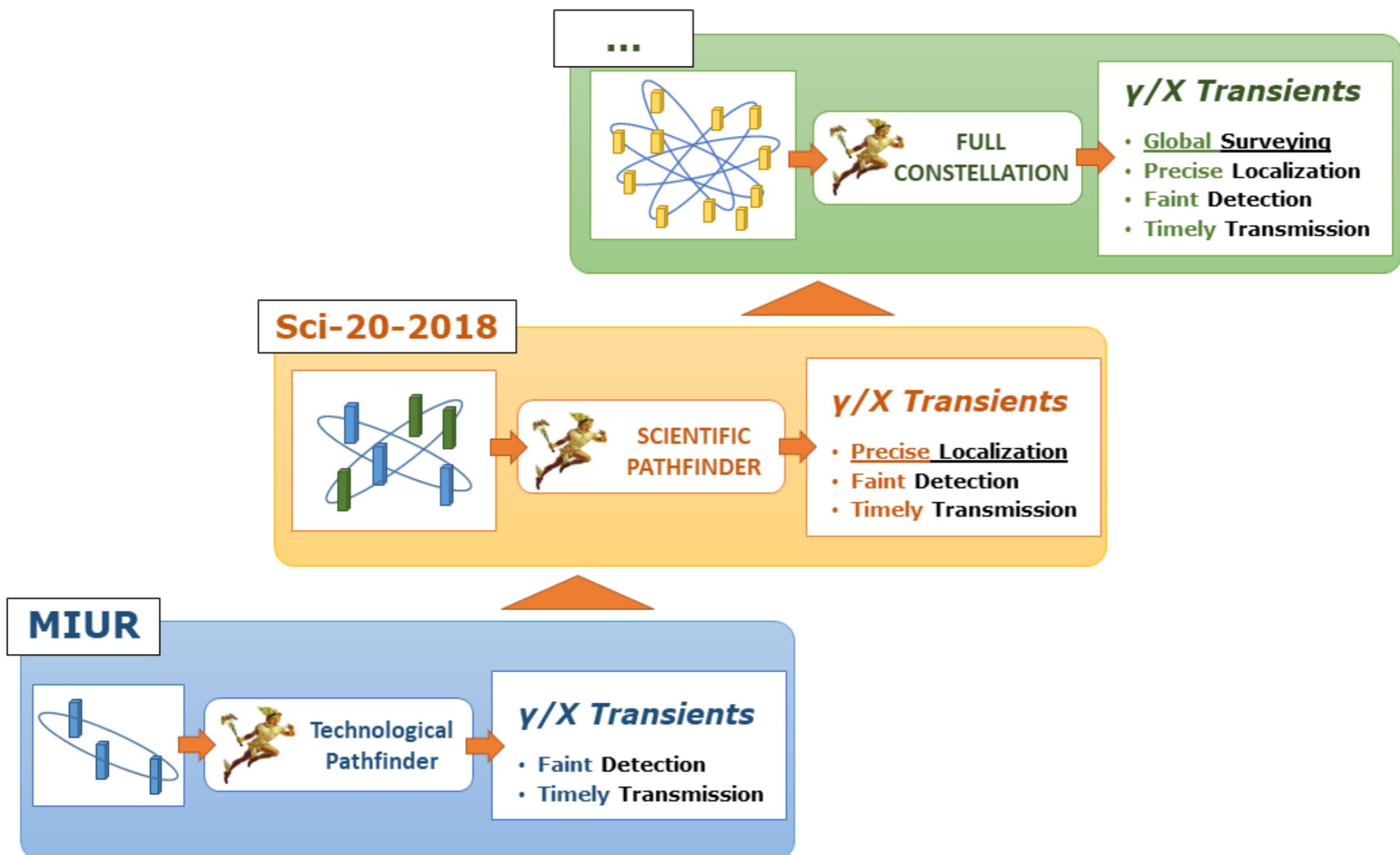
Bright GRBs with msec structure

$\sigma_{\text{Pos(pathfinder)}} \sim 2.4 \text{ deg}$ if $\sigma_{\text{CCF}}, \sigma_{\text{sys}} \sim 0.001 \text{ s}$

$\sigma_{\text{Pos(FC)}} \sim 3 \text{ arcmin}$ if $\sigma_{\text{CCF}}, \sigma_{\text{sys}} \sim 0.001 \text{ s}$

Short GRBs without substructure, risetime fraction of second.

Why HERMES now



- Trend in cost reduction of manufacturing and launching QS

Programmatics 1

Progetto Premiale 2015: **HERMES-Techonogic Pathfinder**

Main objectives:

1. Detect GRBs with simple payload hosted by a 3U CubeSat
 2. Study statistical and systematic errors in the determination of the CCF
- KO May 2018
 - CDR+QR T0+15 QM—> PFM1
 - AR T0+24 —> PFM2+PFM3
 - Launch mid-end 2020 ASI provided
(VegaC maiden flight or Vega, or other opportunities)

Programmatics 2

H2020 SPACE-SCI-20: **HERMES-Scientific Pathfinder**

- Main objectives:
 1. First GRB localization experiment with ≥ 4 CubeSat
 2. Study the systematics associated to the localization
- KO November 2018
- CDR+QR T0+15 QM \rightarrow PFM1
- AR T0+24 \rightarrow PFM2+PFM3
- Launch 2021 (ASI provided)

Programmatics 3

ASI 2019: **HERMES - Advanced Scientific Pathfinder**

- Main objectives:
 1. Nearly all sky coverage
 2. First accurate GRB localization experiment with ≥ 6 CubeSat
- Submitted to ASI September 2018
- Launch 2022? (ASI provided)

HERMES Institutes

- INAF, ASI, PoliMi, UniCagliari, UniPalermo, UniUdine, UniTrieste, UniPavia, UniFedericoll, UniFerrara, FBK, FPM

HERMES is open to ideas and collaboration

Want to be involved? Send an e-mail

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