


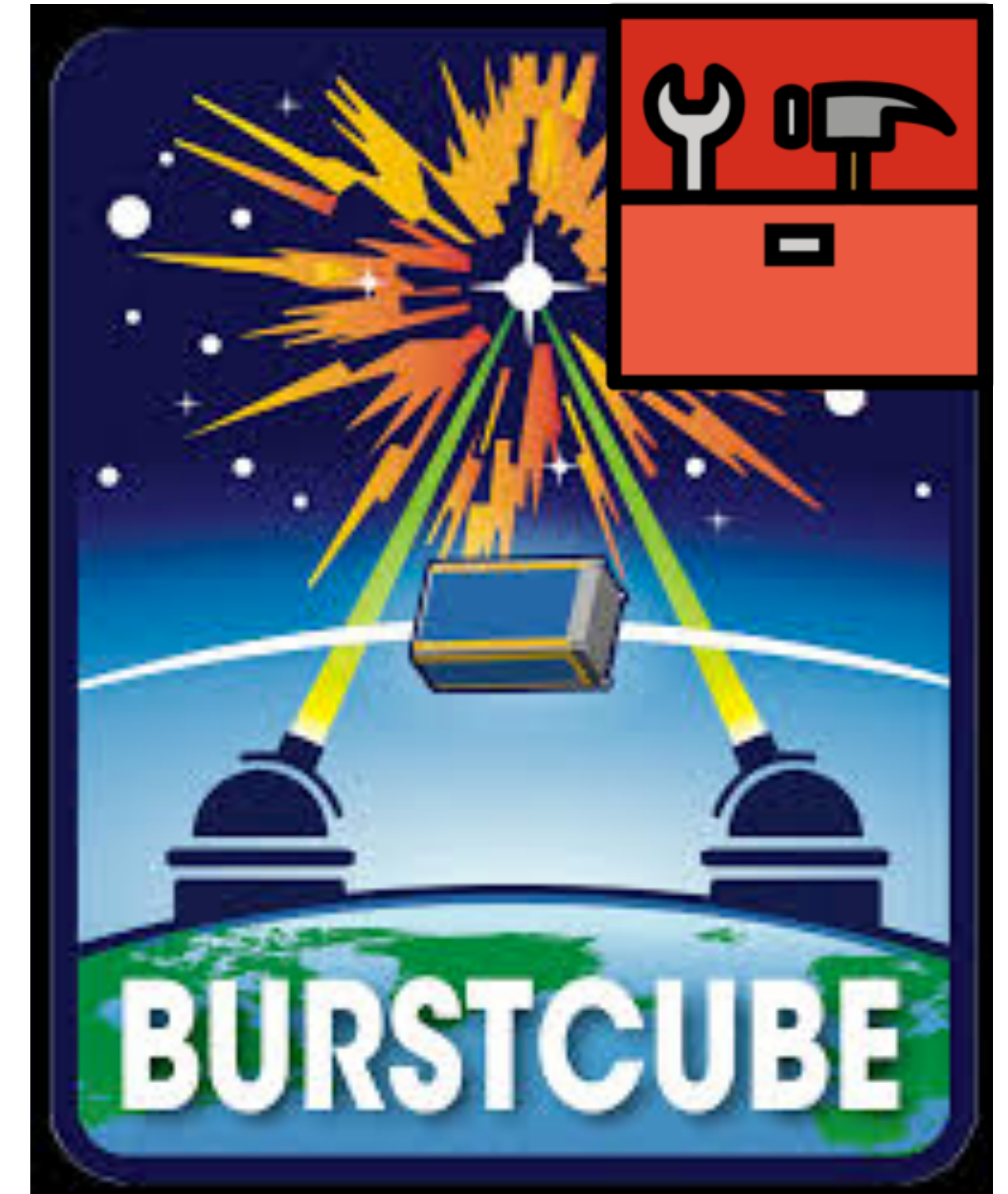
bc-tools



Israel - July 2nd, 2020 - GRB nanosats

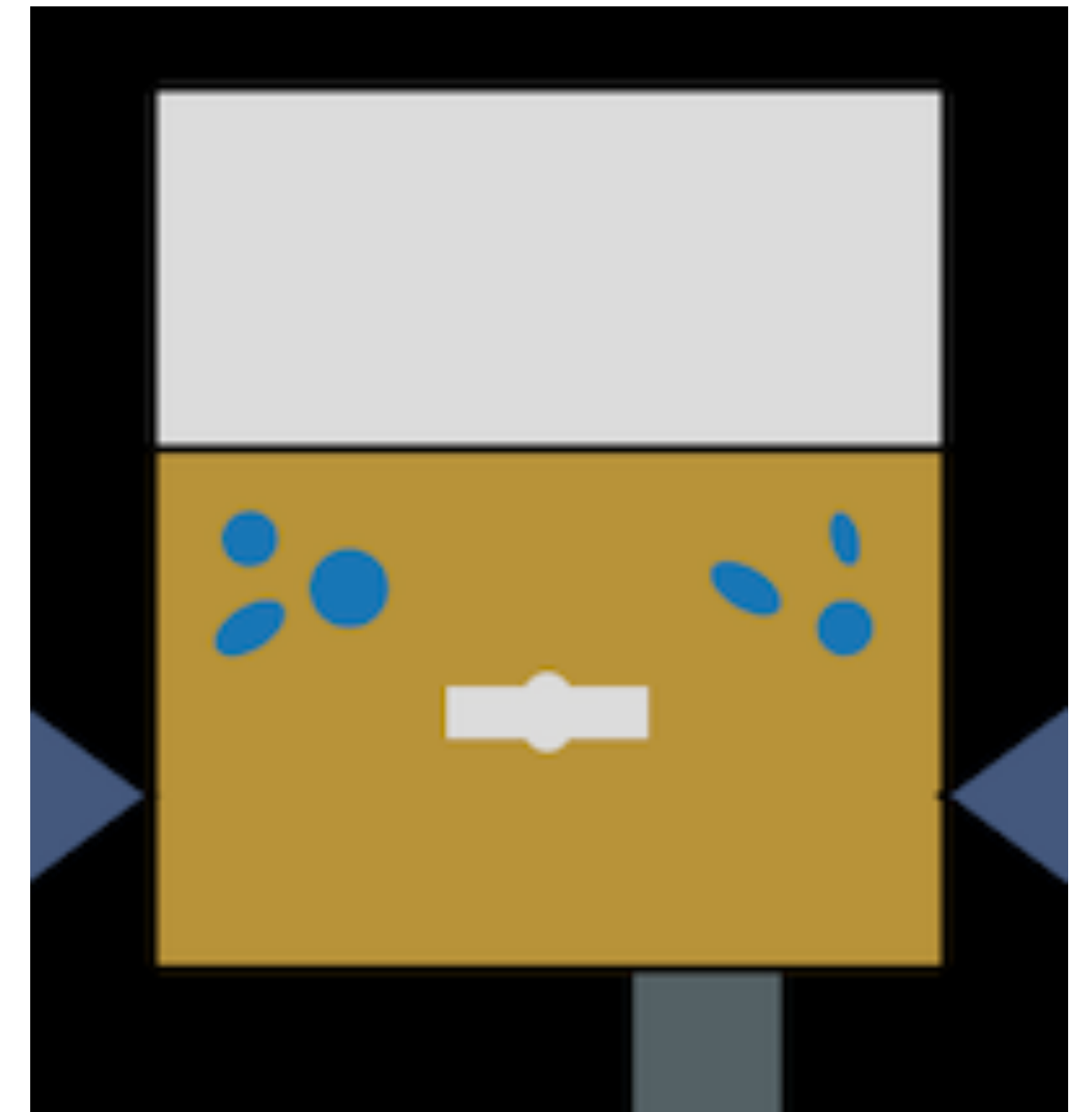
Overview

- bc-tools is BurstCube's main software package
 - Simulations
 - Analysis
- Written in Python
- Currently under development 
- Built around and compatible with gbm-data-tools
- bc-tools is detector-agnostic
 - No hardcoded values
 - Easily adapted for other detectors through a configuration file



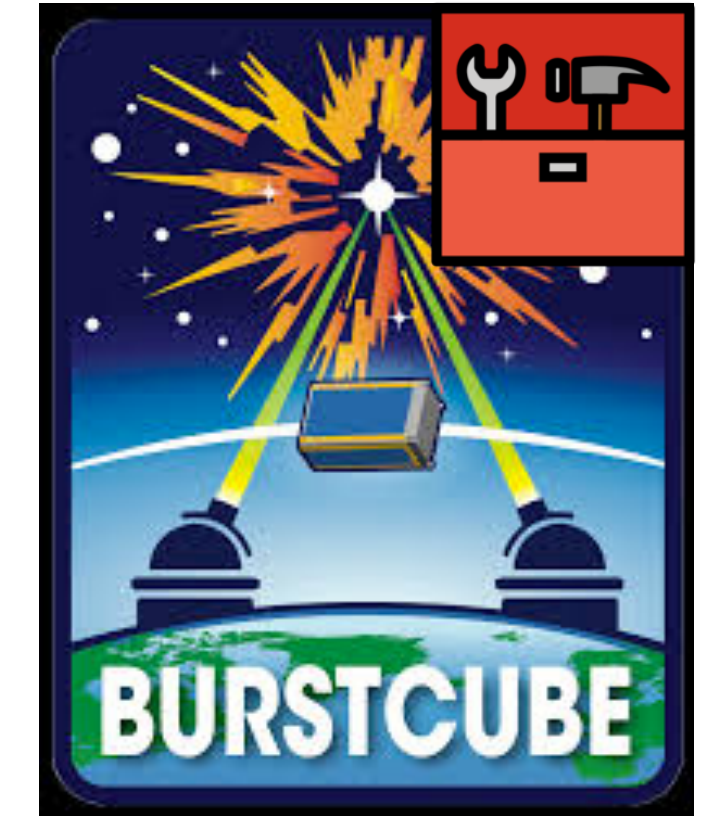
About gbm-data-tools

- Python library developed to analyze GBM data
- Written by Adam Goldstein, William H. Cleveland and Daniel Kocevski
- Can perform most of the tasks we want in a scintillator-based gamma-ray detector:
 - Data binning and light curve generation
 - Background estimation
 - Spectral fitting
 - Source injection
- It has a great high-level API, but also well-documented access to low-level classes
- Built around existing GBM's workflow and data files.
For each burst you have:
 - Data files with counts, such as Time-Tagged Events (TTE)
 - Detector response matrix specific for this event.
One RSP file per detector, which you have to pair manually.

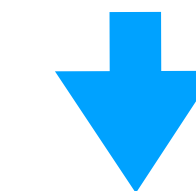


bc-tools as a detector response generator

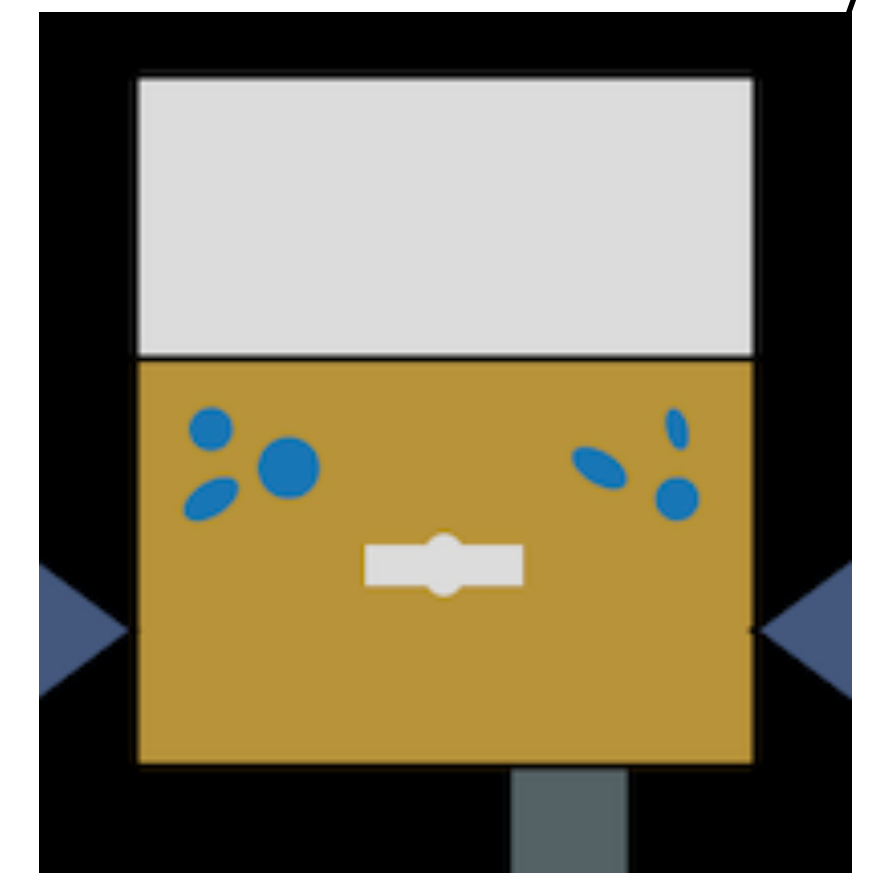
- gbm-data-tools already does most of what we want, we plan to use it and not duplicate efforts
- BurstCube's data files are going to have the same format as GBM (FITS files)
- At first order, all we need is to generate a detector response file
 - The RSP file (also a FITS file) contains the effective area and migration matrix: real energy vs energy channels
 - Corresponds to a specific direction and to a [single] detector
 - Can contain responses for multiple time intervals (e.g. long GRB, spacecraft rocking)
- The app bc-rsp does precisely this:
 - Uses MEGALib for a particle-by-particle MC simulation
 - Build the detector response matrix (a simple 2D array)
 - This is used to construct a GBM's RSP instantiation using the method `from_arrays()`
 - Uses `RSP.write()` to generate a file ready to be used by GBM



```
from gbm.data import RSP
```

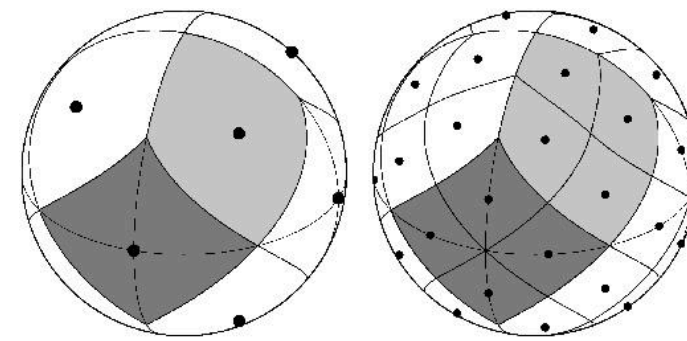


.rsp



The bc-tools full detector response file

- bc-rsp not just generates a .rsp for a given direction and detector, but a full response describing the instrument as a whole
 - All detectors are included, bc-tools knows how to handle them based on their name
 - The response is computed for multiple directions in the sphere:
 - A HEALPix grid is used
 - Response for arbitrary locations are obtained through interpolation.
- This full detector response is saved into a HDF5 file (~GB size)
 - Supports partial loading, only the needed bytes are loaded into memory
- You can extract GBM's RSP objects from there and switch to using gbm-data-tools if you want:



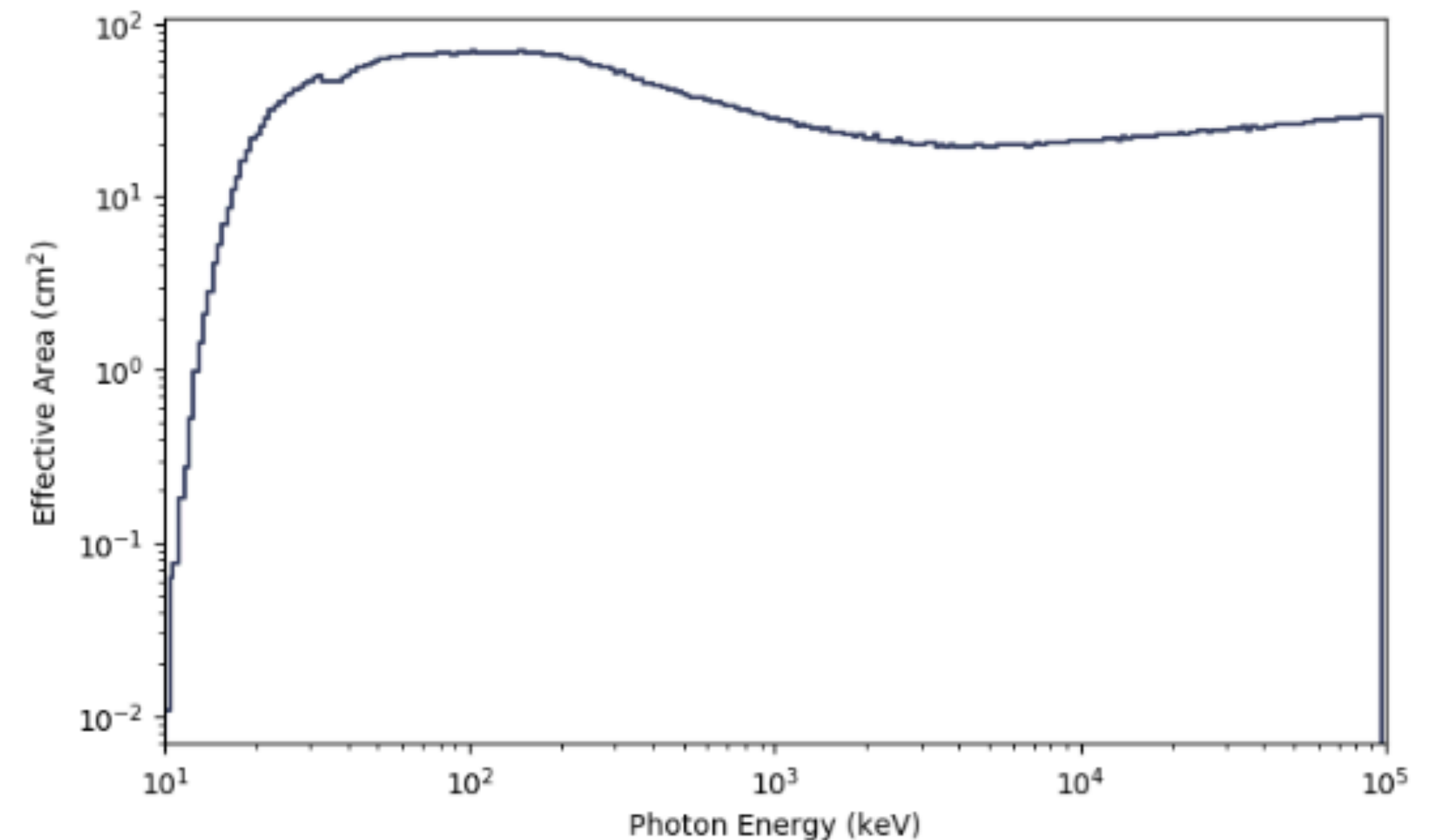
```
from BurstCube.io import FullRSP
from BurstCube.util.coords import SpacecraftCoords
import BurstCube.util.units as u

with FullRSP(data.path/'sims/drm.h5') as frsp:

    rsp = frsp.get_rsp(detector = "SQD0",
                      coords = SpacecraftCoords(45*u.deg, 45*u.deg))

from gbm.plot import PhotonEffectiveArea

effarea_energy = PhotonEffectiveArea(data=rsp)
```



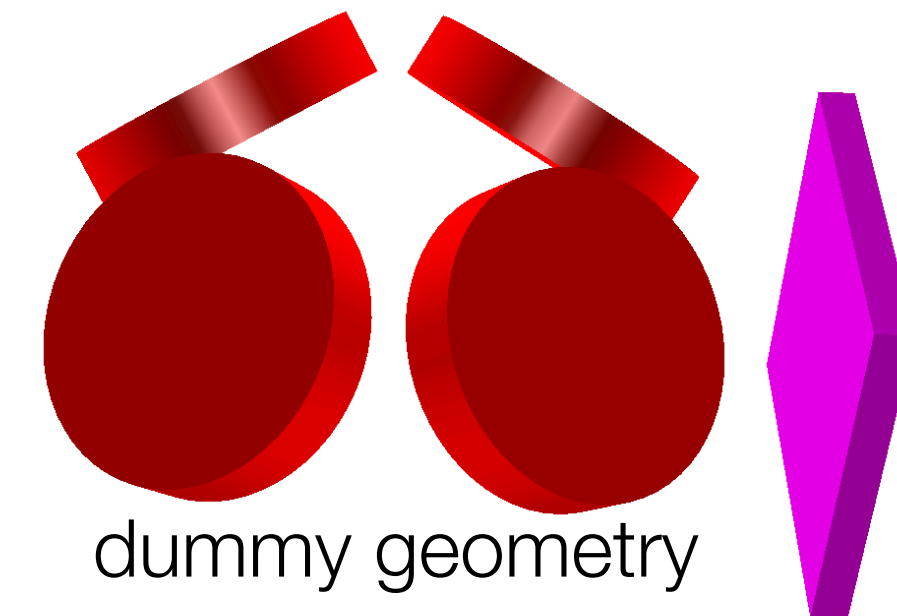
Using bc-tools on its own

- In the final design, gbm-data-tools will be under the hood
- The user will have access to the full detector response:
 - Timing and detector matching will be done automatically
 - The full response will be used for localization as well (not part of gbm-data-tools)
- Will output the results from the maximum likelihood calculations. This allows to:
 - Perform a more in-depth analysis. e.g. localization vs spectrum correlations
 - Combine the data easily with other experiments:
 - LIGO-Virgo joint analysis
 - 3ML plug-in
- The ability to get a GBM-compatible .rsp will always be there
 - Backward compatibility with software like *Xspec*



Build your own instrument

- An instrument is defined by:
 - A Geomega (MEGALib) geometry
 - Names for the single detectors
 - Calibrated instrument effects. e.g.
 - Energy resolution vs energy
 - Efficiency vs energy
- A YAML config file holds these and most other parameters
 - Job specific (e.g. random seeds) are passed through command-line options
 - Any parameter can be modified on the fly from the command-line for easy testing. e.g.
`--override simulations:geometry=calib_bc.geo`

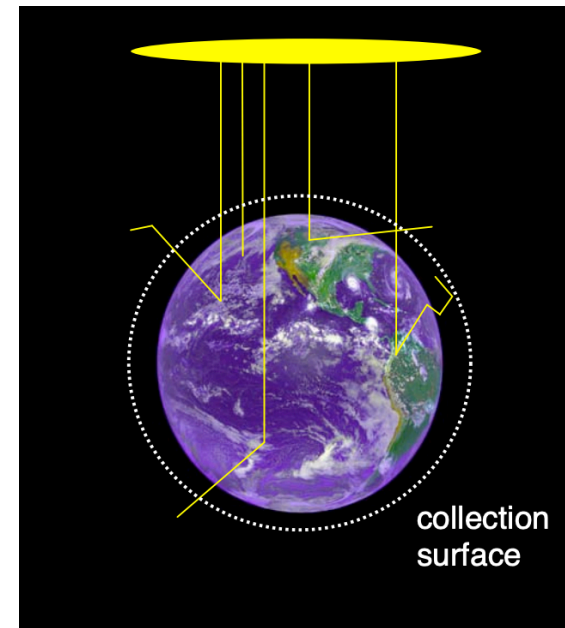


```
1 # Example of a configuration file for BurstCube
2
3 # Note: Since some parameters are long lists, this is easier to
4 # read if you deactivate line wrapping e.g. less -S bc_config.yaml
5
6 # Stuff related to MEGALib simulation
7 simulations:
8   #Geomega file, relative to this file
9   geometry: "geometry/BurstCube.geo.setup"
10
11 # When a calorimeter is hit Cosima reports the location of the
12 # center of mass. There is no ID or detector name in the .sim file.
13 # This location is not trivial to compute from the geometry file, the
14 # simplest way to get this locations is to run Cosima once and
15 # copy-paste them here. e.g. for SDQ0 you get a hit entry like this:
16 # HTsim 4; 4.88500; 4.88500; -0.03536; 120.12985;0.00000e+00;2
17 # This makes me sad :(
18 detectors:
19   location: {"4.88500, 4.88500, -0.03536": "SQD0",
20             "-4.88500, 4.88500, -0.03536": "SQD1",
21             "-4.88500, -4.88500, -0.03536": "SQD2",
22             "4.88500, -4.88500, -0.03536": "SQD3"}
23
24 # Spectrum of thrown particle. As long as we have narrow energy bin,
25 # the result do not strongly depend on this
26 # See Spectrum class for options
27 # Note that the normalization is arbitrary.
28 spectrum:
29   name: "PowerLaw"
30   args:
31     norm: 1
32     index: 1
```


Final remarks

- Generation of full detector response files is pretty much done
- Source localization code is the next step
- Automatic science pipelines still need to be developed
- A possible challenge: atmospheric scattering

- The atmospheric simulation itself will not be part of bc-tools
- We are hoping GBM's simulations are good enough for BurstCube.



- The project is public and we welcome contributions through merge requests:
gitlab.com/burstcube/bc-tools

